# The dimensional evolution of structure and dynamics in hard sphere liquids

Cite as: J. Chem. Phys. **156**, 134502 (2022); https://doi.org/10.1063/5.0080805 Submitted: 03 December 2021 • Accepted: 05 February 2022 • Published Online: 04 April 2022

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# The dimensional evolution of structure and dynamics in hard sphere liquids

Cite as: J. Chem. Phys. 156, 134502 (2022); doi: 10.1063/5.0080805 Submitted: 3 December 2021 • Accepted: 5 February 2022 • Published Online: 4 April 2022 • Publisher Error Corrected: 07 April 2022







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Note: This paper is part of the JCP Special Topic on Slow Dynamics.

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#### **ABSTRACT**

The formulation of the mean-field infinite-dimensional solution of hard sphere glasses is a significant milestone for theoretical physics. How relevant this description might be for understanding low-dimensional glass-forming liquids, however, remains unclear. These liquids indeed exhibit a complex interplay between structure and dynamics, and the importance of this interplay might only slowly diminish as dimension d increases. A careful numerical assessment of the matter has long been hindered by the exponential increase in computational costs with d. By revisiting a once common simulation technique involving the use of periodic boundary conditions modeled on  $D_d$  lattices, we here partly sidestep this difficulty, thus allowing the study of hard sphere liquids up to d = 13. Parallel efforts by Mangeat and Zamponi [Phys. Rev. E 93, 012609 (2016)] have expanded the mean-field description of glasses to finite d by leveraging the standard liquid–state theory and, thus, help bridge the gap from the other direction. The relatively smooth evolution of both the structure and dynamics across the d gap allows us to relate the two approaches and to identify some of the missing features that a finite-d theory of glasses might hope to include to achieve near quantitative agreement.

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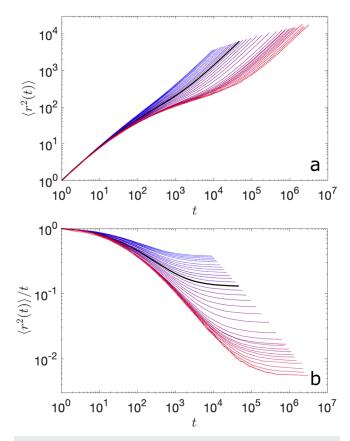
# I. INTRODUCTION

Hard spheres are a classical minimal model for the structure and dynamics of real (yet simple) liquids. 1,2 The model captures many of the salient features of liquids while being simple enough to treat using both analytical theory and computer simulations. From a theoretical standpoint, in particular, there have been two main thrusts: (i) developing models that capture the features of lowdimensional (d = 2 and d = 3) fluids reasonably well and (ii) developing a finite-d mean-field description that becomes exact in the limit  $d \to \infty$ . Both approaches have been met with success,<sup>1</sup> but the extrapolation of the former to high densities and of either to intermediate dimensions (e.g.,  $4 \le d < 20$ ) leads to various quantitative inconsistencies that may reflect weaknesses in our physical understanding of these systems.

In order to expand on this point, consider first the high-density regime (see Fig. 1). A low volume fraction ( $\varphi$ ) Brownian fluid of hard spheres (HS) begins as a Fickian fluid characterized by a purely diffusive mean squared displacement (MSD). Upon compression,

non-Fickian diffusion first emerges at a dynamical onset,  $\varphi_{nf}$ . Upon further compression, liquid dynamics turns increasingly sluggish as particles first become transiently caged by each other and then become nearly arrested. In mean-field descriptions, this sustained caging is associated with a topological change in the free energy landscape at the dynamical (or mode-coupling theory) transition,  $\varphi_d$ . Although this transition is formally avoided in finite-d systems, the underlying crossover has long been extracted from the pseudo-divergence of the structural correlation time observed in high density fluids. 4,10,1

Signatures of both  $\varphi_{\rm d}$  and  $\varphi_{\rm nf}$  have long been sought out in the intricate structure of low-dimensional liquids, but  $d \to \infty$ fluids contain no such feature. For hard spheres, the pair correlation function is then flat beyond contact, and all higherorder correlations are trivially factorizable.2 The structure of lowdimensional liquids, which is expected to smoothly evolve toward the high-dimensional description, must therefore at least quantitatively perturb the mean-field predictions for these two signature features.



**FIG. 1.** (a) MSD of a hard sphere fluid in d=5 averaged over 1000 realizations. The system smoothly goes from Fickian at  $\widehat{\varphi}=1$  (blue line), through the onset of non-Fickian diffusion at  $\widehat{\varphi}_{\rm nf}=1.36$  (black line), and up to a dense fluid  $\widehat{\varphi}=1.648$  (red line) near  $\widehat{\varphi}_{\rm d}\approx 1.701$ . (b) The diffusion constant D can be extracted from the long-time plateau of  $\langle r(t) \rangle/t$  after rescaling the short-time dynamics. Fits to Eq. (13) are overlaid.

The chasm between what has been extracted from existing numerical simulations and the  $d \to \infty$  solution is, however, too vast to say for sure whether this proposal holds. For instance, meanfield estimates for the dynamical transition computed using both the Percus-Yevick (PY) and Hypernetted Chain (HNC) approximations in  $d = 2-70^{12}$  deviate substantially from each other and from simulation estimates. 13,14 While both the HNC and PY approximations capture the  $d \to \infty$  value of  $\varphi_d$ , there is no indication that this high-dimensional approach is the same as that of HS or, thus that the gap between simulation and these two theories tends to vanish with increasing d. The observed discrepancy might not only partly result from the small system sizes previously used in simulations but might also attest to the structural inadequacies of the liquid-state descriptions. Given that our understanding of glasses and supercooled liquids has since developed, can the dimensional gap now be properly bridged? Hardware improvements over the last decade offer some hope on the computational side but remain far from sufficient. Fortunately, algorithmic techniques have also since been refined. In particular, the development of smart boundary conditions for high- d systems has significantly extended the numerical reach.  $^{15}$ 

These advances here allow us to obtain a near quantitative agreement between theory and numerics in some respects. More specifically, in this work, we show that the relevant features of hard sphere liquids converge to those of HNC fluids as d increases and that for  $d \ge 13$ , they become almost indistinguishable, thus possibly making the study of higher-dimensional fluids essentially unnecessary. We also show that the HNC-based predictions for  $\varphi_A$  tend to track the numerical results but remain quantitatively distinct. The rest of this paper is structured as follows. Section II details the numerical techniques that allow us to equilibrate dense liquids up to d = 13. Section III describes the evolution of fluid pair structure with dimension, Sec. IV describes the evolution of dynamics with dimension, and Sec. V concludes by briefly considering how further theoretical advances might strengthen the quantitative interplay between the mean-field theory of glasses and numerical simulations.

# **II. MODEL AND METHODS**

Throughout this work, we consider hard sphere liquids composed of d-dimensional hyperspheres of diameter  $\sigma$  with  $4 \le d$  $\leq$  13. For  $d \geq$  4, monodisperse spheres suffice because crystallization, while thermodynamically possible, 15,16 is then effectively always Initial states are first prepared using Poissonsuppressed.17 distributed soft hyperspheres whose energy is minimized to zero, thus yielding a valid hard sphere configuration. These configurations are then equilibrated via standard Monte Carlo (MC) dynamics, 21 using attempted individual particle step sizes of  $\kappa\delta\sigma$ along each dimension with  $\kappa \in [-1,1)$  distributed uniformly at random and  $\delta \approx 1/(6d)$  chosen so as to optimize the equilibration time at high  $\varphi$ . The numerical results are, however, rather insensitive to this specific choice, given the rescaling described in Sec. IV B. In the rest of this section, we describe various aspects of the numerical implementation, paying particular attention to features that enable the efficient consideration of higher-dimensional systems, such as the simulation box shape and the system size N.

# A. Equilibration and sampling

The timescale for structural decorrelation (and hence equilibration) is determined from the characteristic decay of the particle-scale self-overlap,

$$Q(t) = \frac{1}{N} \sum_{i=1}^{N} \Theta(a\sigma - |\mathbf{r}_i(t) - \mathbf{r}_i(0)|), \tag{1}$$

where  $r_i(t)$  is the position of particle i at time t,  $\Theta$  is the Heaviside function, and a is the length on the scale of the typical particle cage size around  $\varphi_d$ . (In this work, a=0.3. Although asymptotic 1/d corrections to a are expected, in finite-d, the dependence has been found to be relatively weak.<sup>22</sup> In any event, taking a constant value merely overestimates the relaxation time  $\tau$ , which has no bearing on the subsequent analysis, as further discussed in Sec. IV B.) In other words, Q(t) captures the fraction of particles having moved a distance larger than  $a\sigma$  at time t. This function is known to generically decay as a stretched exponential  $Q(t) \sim \exp(-(t/\tau)^{\zeta})$  with  $\zeta = 1$  for

Fickian fluids and to decrease as  $\varphi_d$  is approached. The characteristic relaxation time  $\tau$  is then implicitly defined as  $Q(\tau) = 1/e$ .

Lower density systems are equilibrated for at least  $3\tau$  before recording the pair correlation function, g(r), and the mean squared displacement (MSD),  $\langle r^2(t) \rangle$ . This starting configuration is also used for generating additional equilibrated configurations, whereupon 1000 realizations are averaged. To properly characterize higher density configurations, and, in particular, to extract  $\varphi_d$ , simulations need to be run for much longer upon its approach. An equilibration of at least  $30\tau$  and as few as ten independent realizations are then used.

#### B. Structure and pressure

For distances r less than half the box size, the standard scheme to extract g(r) is used (see, e.g., Ref. 21, Sec. 4.4) because the shell of neighbors is then perfectly spherical. For r larger than half the box size, the shell encompassing r intersects the periodic box limits. Although both the particle count and the shell volume are smaller than in the perfect spherical shell, the sphere density is unaffected and, thus, relevant structural information can still be extracted. Because of the relatively complex geometry of the truncated shell, however, its volume is obtained by a simple Monte Carlo integration rather than analytically.

From the virial theorem, the reduced pressure p for hard spheres can be determined from the contact value of the pair correlation,

$$p(\varphi) = \frac{\beta P}{\rho} = 1 + 2^{d-1} \varphi g(\sigma^+), \tag{2}$$

where  $\beta = 1/k_BT$  is the inverse temperature, P is the standard pressure, and  $\rho = N/V$  is the number density for a simulation box of volume V. In order to compare thermodynamic quantities across dimensions, we correct for their asymptotic  $d \to \infty$  scaling. In particular, the volume fraction is rescaled as  $\widehat{\varphi} = 2^d \varphi/d$  and the reduced pressure as  $\widehat{p} = p/d$ . Equation (2) then simplifies to

$$\widehat{p}(\widehat{\varphi}) = \frac{1}{d} + \frac{\widehat{\varphi}}{2}g(\sigma^+). \tag{3}$$

This rescaling makes clear that given that g(r) is a pure step function in limit  $d \to \infty$ , the equation of state then becomes<sup>23</sup> [see also Ref. 2, Eq. (2.61)]

$$\widehat{p} = \frac{\widehat{\varphi}}{2} \quad \text{for} \quad d \to \infty.$$
 (4)

#### C. Boundary conditions

As mentioned in Sec. II B, the standard scheme for determining the radial distribution function is valid up to the inscribed sphere radius of the periodic box. Beyond this point, a particle may have multiple periodic images of a same neighbor within its spherical shells, which give rise to unphysical structural correlations. In this sense, the choice of periodic boundary conditions is optimized when the inscribed sphere radius is maximal for a given box volume under the periodic tiling constraint of  $\mathbb R$ . Optimizing boundary conditions is thus equivalent to optimizing sphere packing in a given d. The conventional cubic periodic boundary condition, which corresponds

to a simple cubic tiling, is clearly not optimal because this sphere packing lattice is not the densest in any  $d \ge 2$ .

This realization is far from novel. The efficiency of noncubic periodic boxes attracted interest historically when computer resources were limited. Tet, the resulting efficiency improvement was not found to be sufficiently significant in d=3 (\$\approx 40\%), in the context of rapidly accelerating computer hardware, especially given the coding apprehension associated with such an approach. As d increases, however, hypercubic periodic boxes become increasingly inefficient, and non-cubic lattice packings can give rise to more sizable efficiency gains. Realizing that the key implementation for simulating some of these boxes can be straightforwardly adapted from quantizing algorithms in information theory. further motivates their reconsideration. For a common simulation code structure, the only significant algorithmic change concerns the minimal image convention, and the structure of that change is often independent of d.

More specifically, we first construct the periodic box scheme for  $Z_d$  (cubic) and  $D_d$  (checkerboard, such as face-centered cubic in d=3) based lattices. We denote f(x) as the closest integer to x, and likewise  $f(x)=(f(x_1),\ldots,f(x_n))$  for an n-dimensional vector x. Clearly,  $\delta x=x-f(x)$  is the minimum image of x to the origin in (hyper-)cubic periodic boundary conditions,  $\mathbb{Z}^n$ , and computing it requires d operations. We also define  $f_2(x)$  as the second-nearest integer to x, and

$$f_2(\mathbf{x}) = (f(x_1), f(x_2), \dots, f_2(x_k), f(x_{k+1}), \dots, f(x_d)), |x_k - f(x_k)| \ge |x_i - f(x_i)| \ \forall i$$
 (5)

is the second-nearest integer vector to x in  $\mathbb{Z}^d$ . The  $D_d$  minimum image of x is chosen to be

$$\delta(\mathbf{x}) = \begin{cases} \mathbf{x} - f(\mathbf{x}), & \sum_{i=1}^{d} f(x_i) \text{ is even,} \\ \mathbf{x} - f_2(\mathbf{x}), & \sum_{i=1}^{d} f(x_i) \text{ is odd.} \end{cases}$$
(6)

This construction defines the  $D_d$  periodic box, and its calculation requires roughly 4d operations.<sup>28</sup> In addition, the  $D_d$  lattice has a maximum inscribed radius  $R_I = 1/\sqrt{2}$  and a circumradius of either 1 for  $d \le 4$  or  $\sqrt{d}/2$  for d > 4. Therefore, the volume of a  $D_d$  periodic box is 2,<sup>29</sup> and the packing fraction of the  $D_d$  lattice is  $2^{\frac{d-2}{2}}$  times greater than  $\mathbb{Z}^d$ , yielding an efficiency gain of that same factor.

The furthest points to lattice sites in  $D_d$  lattices (known as deep holes) are at  $(\frac{1}{2},\ldots,\frac{1}{2})$  and its periodic images. The distance of the deep hole to a lattice site,  $\sqrt{d}/2$ , becomes no less than twice of the inscribed radius for  $d \geq 8$ . Therefore, one can then slide a second  $D_d$  lattice within these holes without changing the inscribed radius of the Voronoi tessellation. The volume of a periodic box is then cut in half. This construction results in  $D_d^+$  lattices whose symmetry is only valid in even dimensions. The minimum image of  $\boldsymbol{x}$  in the  $D_d^+$  lattice is either  $\delta \boldsymbol{x}$  or  $\delta(\boldsymbol{x}-\frac{1}{2})+\frac{1}{2}$ , whichever has the smaller norm. In particular, taking d=8 gives an  $E_8$  periodic box and also corresponds to the densest sphere packing in that dimension. For odd dimensions in  $d \geq 9$ , one can also slide another copy of  $D_d$ , but now by an offset of  $\boldsymbol{a}=(\frac{1}{2},\ldots,\frac{1}{2},a_d)$ , where  $a_d \in \mathbb{R}$  is an arbitrary real number. By convention, taking  $a_d=0$  gives the lattice  $D_d^{0+}$ . The minimum image of  $\boldsymbol{x}$  in  $D_d^{0+}$  lattices is either  $\delta \boldsymbol{x}$  or  $\delta(\boldsymbol{x}-\boldsymbol{a})+\boldsymbol{a}$ , whichever has the smaller

norm. In particular,  $D_9^{0+}$  is a realization of the  $\Lambda_9$  lamellar lattice, which is the densest known sphere packing structure in d = 9.

The inscribed radius of both  $D_d^+$  and  $D_d^{0+}$  periodic boxes remains  $1/\sqrt{2}$ , and their volume is 1—half that of  $D_d$  boxes. The implementation of  $D_d^+$  and  $D_d^{0+}$  periodic boxes thus doubles the efficiency over  $D_d$ . Because it involves computing two sets of  $D_d$ minimum images and squared distances, the minimum image algorithm takes 13d operations in total. (For the  $E_8$  lattice, a special quantization algorithm takes only 72 operations.<sup>28</sup>) In practice, we observe that the minimum image determination is roughly three times as computationally demanding as for  $D_d$ . The efficiency gain of these lattices is thus canceled if the minimum image determination is the computational bottleneck, as is the case in liquid simulations. (It may be possible to do better in certain higher dimensions, such as for the remarkably dense Leech lattice,  $\Lambda_{24}$ , which is a factor of  $2^{13}$ denser than  $D_{24}$ , and for which a fast quantization algorithm takes only 55 968 steps.<sup>28</sup> A factor of roughly 10—compared to  $4d \times 2^{13}$ = 786 432—should thus be gained from this geometry.) In practice, we, thus use simulation boxes with standard cubic periodic boundaries in d = 3 and d = 4 and  $D_d$  periodic boxes for  $d \ge 5$ .

# D. System size considerations

A particularly subtle issue in higher dimensional simulations is the system size. Careful consideration has thus been given to this matter in prior simulation studies. <sup>13,30</sup> For the sake of the current work, it suffices to recall that for proper thermodynamic quantities to be extracted, systems should be much larger than the largest correlation length. Once that target is attained, remaining corrections should scale as 1/N. Hence, as long as the systems simulated are larger than the static and dynamical correlation lengths (hydrodynamic correlations are not relevant in MC simulations because momentum is not conserved), then only trivial system size corrections should persist.

In low-dimensional dense fluids, various static features have been argued to play a key dynamical role. The slow convergence of the virial series further suggests that multi-particle correlations also then contribute significantly to the equation of state. As d increases, however, structural correlations are expected to steadily vanish and structural anomalies that couple to dynamics are expected to homogenize, at least for  $\widehat{\varphi} < \widehat{\varphi}_d$ . Although the dynamical correlation length is expected to diverge at  $\widehat{\varphi}_d$ —as the crossover hardens with increasing d—the critical regime is expected to be fairly narrow. <sup>8,13</sup>

Therefore, as long as we avoid introducing correlations by, say, having a particle interact with itself through the periodic boundary, then relatively modest simulation box sizes should suffice. For these purposes, it is useful to note that the shortest distance across the simulation box is  $2R_I$ , where  $R_I$  is the maximum inscribed radius of the simulation box (Table I). Writing the system size in terms of the number of spheres that can fit end-to-end,  $n_\sigma$ , gives  $2R_I = n_\sigma \sigma$  and, thus,

$$N = \frac{\widehat{\varphi}dV\Gamma\left(\frac{d}{2} + 1\right)}{\pi^{d/2}} \left(\frac{n_{\sigma}}{2R_{I}}\right)^{d}.$$
 (7)

The limited role played by structural and dynamical correlations can separately be validated by verifying that the properties of the systems studied converge smoothly to those of the  $d \to \infty$  fluid, which are known exactly.

TABLE I. Summary of select periodic box properties.

Symmetry	Valid in	Volume (V)	Inscribed radius $(R_I)$	Packing efficiency w.r.t. $\mathbb{Z}^d$
$\mathbb{Z}_d$	$d \ge 1$	1	$\frac{1}{2}$	1
$D_d$	$d \ge 2$	2	$1/\sqrt{2}$	$2^{\frac{d-2}{2}}$
$D_d^+$	Even $d \ge 8$	1	$1/\sqrt{2}$	$2^{rac{d}{2}}$
$D_d^{0+}$	Odd $d \ge 9$	1	$1/\sqrt{2}$	$2^{rac{d}{2}}$
$\Lambda_{24}$	d = 24	1	1	$2^{24}$

In practice, we find that system sizes such that  $n_{\sigma} \ge 2.3$  at  $\widehat{\varphi}_{\rm d}$  suffice. For  $d \ge 15$ , however, this criterion would require  $N \gg 10^5$ , which lies beyond the reach of our computational resources.

# III. DIMENSIONAL EVOLUTION OF FLUID STRUCTURE

In this section, we compare the dimensional evolution of structural observables with theoretical predictions. We specifically consider the fluid equation of state and the shell structure of g(r) (Fig. 1).

#### A. Equation of state

The liquid equation of state is one of the simplest observables to evaluate. Its relative crudeness is compensated by a dearth of available theoretical predictions. Here, we specifically compare numerical results [using Eq. (3)] with virial, PY, and HNC<sup>33</sup> predictions. For further analysis, these results are succinctly fitted to a generalized Carnahan–Starling (CS) form <sup>13,31,34</sup>

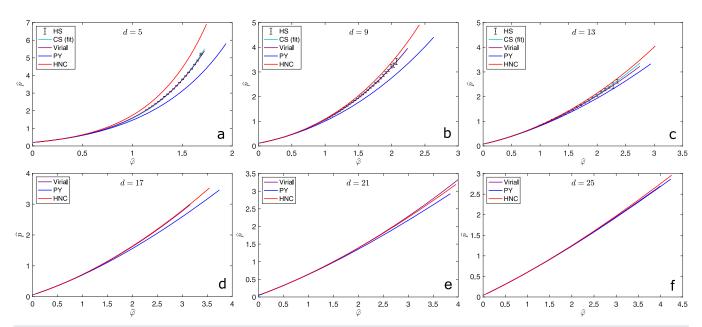
$$\widehat{p} = \frac{1}{d} + \frac{\widehat{\varphi}}{2} \left[ \frac{1 + \widehat{A}\widehat{\varphi}}{(1 - \varphi)^d} \right], \tag{8}$$

which has only one free parameter,  $\widehat{A}$  (Fig. 2). This form captures numerical results quite well over the regime of interest,  $\widehat{\varphi}_{nf} < \widehat{\varphi} < \widehat{\varphi}_{d}$ . (Higher-order corrections would mostly account for deviations at lower  $\widehat{\varphi}^{35}$  and hence are not considered.) As expected, we find that  $\widehat{A} \to 0$  as  $d \to \infty$ , thus recovering Eq. (4). It is interesting to compare this fitted factor with the generalized CS treatment of Song *et al.*, <sup>31</sup> for which

$$\widehat{A}^{CS} = \frac{d}{2} \left( \frac{B_3}{B_2^2} \right) - \frac{d^2}{2^d} = d \left[ 1 - \frac{{}_2F_1\left(\frac{1}{2}; \frac{1-d}{2}; \frac{3}{2}; \frac{1}{4}\right)}{\beta\left(\frac{1}{2}, \frac{1+d}{2}\right)} \right] - \frac{d^2}{2^d}$$

$$= dI_{\frac{3}{4}} \left( \frac{d+1}{2}, \frac{1}{2} \right) - \frac{d^2}{2^d} \sim \sqrt{\frac{6d}{\pi}} \left( \frac{3}{4} \right)^{d/2} \left[ 1 - \frac{15}{4d} + \mathcal{O}\left(\frac{1}{d^2}\right) \right] - \frac{d^2}{2^d}, \tag{9}$$

where  $B_n$  is the virial coefficient of order n,  ${}_2F_1$  is the hypergeometric function,  $\beta$  is the beta function, and  $I_x(a,b)$  is the regularized incomplete beta function. A high-d asymptotic scaling form is also provided. [Although this form could be developed to higher order for marginally better asymptotic scaling,  $^{36}$  the corrections would nevertheless become increasingly significant for d < 25 because the  $I_x(a,b)$  scaling only converges for  $x \in [0,\frac{1}{2})$ , even though an asymptotic expression can be obtained for  $x \in (0,1)$ .] This scaling form



**FIG. 2.** Liquid equations of state for (a) d=5, (b) d=9, (c) d=13, (d) d=17, (e) d=21, and (f) d=25. The first three panels compare the simulation results (HS) with the fitted Carnahan–Starling form (CS), the Padé resummed virial series, and the Percus–Yevick (PY) and Hypernetted Chain (HNC) predictions. Only the latter three quantities appear in the last three panels. From these results, it is clear that both the CS form and the resummed virial series remain good descriptors of the equation of state in that  $\widehat{\varphi}$  regime in high d. It is also clear that the equation of state is systematically better described by the HNC than by the PY approximation as d increases. Remarkably, for  $d \ge 13$ , the HNC predictions have essentially converged with the simulation results.

might give the impression that deviations (on a linear scale) from the  $d \to \infty$  results appear markedly more significant for  $d \lesssim 50$ , but the convergence is actually smooth (see also Sec. IV B). Simulation estimates for  $\widehat{A}$  from HS simulations (for  $4 \le d \le 11$ ), the virial (for  $13 \le d \le 50$ , see below), and the HNC lend further support to this scaling form. The liquid–state approach to the  $d \to \infty$  description is therefore highly non-trivial.

We next consider the reliability of the virial series. For hard spheres, coefficients have been reported up to n = 10 in  $4 \le d \le 8^{37}$  and up to n = 30 for  $d \ge 9$ . In order to make use of that series at high densities, however, it is necessary to resum it to capture its (apparent) divergence in the liquid phase. As is standard,  $^{15,16,37-40}$  we use  $[\ell, m]$  Padé approximants, which given the number of known terms in the virial series  $n = \ell + m + 1$  are

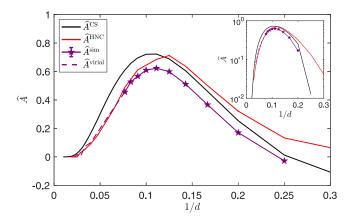
$$p = \frac{1 + \sum_{i=1}^{\ell} b_i \rho^i}{\sum_{i=1}^{m} \bar{b}_i \rho^i},$$
 (10)

where  $b_i$  and  $\bar{b}_i$  are determined from the set of coefficients  $\{B_1,\ldots,B_n\}^{41}$  There is no *a priori* correct choice for  $\ell$ , m, and n; Padé approximants are notoriously ill behaved, as evidenced by the emergence of spurious poles below  $\widehat{\varphi}_d$ , for even small changes to  $\ell$  and m. Nevertheless, conventional wisdom suggests that the best fits occur for  $\ell \approx m.^{15,16,37-40}$  We, thus, here use  $m-1 \le \ell \le m+1$  and choose the lowest order  $\ell$  for which no spurious pole appears. Within these constraints, we find the [4,5] approximant in  $d \le 8$  and either the  $[\ell,\ell]$  or the  $[\ell,\ell-1]$  approximant in d > 8 to most reliably converge. For  $d \le 13$ , a solid agreement with numerical results

is then obtained. It is thus reasonable to extract  $\widehat{A}$  from the virial series up to d = 30. For d > 30, however, the relative error on the series coefficients makes this process numerically unreliable.

Having validated the numerical and resummed virial results, we can now consider the accuracy of the HNC and PY predictions. Both schemes properly converge to the  $d \to \infty$  solution,<sup>23</sup> but their asymptotic scaling differs, and it is unclear how well either describes high but finite-d systems. We note that although the HNC results are fairly crude in low dimensions, they recapitulate the numerical data increasingly well as d increases. For d = 13, the results are already pretty close, and for d = 17, the virial and the HNC predictions overlap within the uncertainty on the resummation scheme [Fig. 2(d)]. The expectation that the HNC approaches the equation of state asymptotically well as d increases 12,42 is therefore validated. By contrast, PY predictions remain quite far off the mark for d < 25. They systematically undershoot both numerical and virial results. Hence, although PY does a remarkable (and likely fortuitous) job in d = 3 and is correct in  $d \to \infty$ , its description of intermediate dimensions is inadequate. It cannot be relied upon for estimating other observables.

The convergence of the HNC prediction onto the virial results as d increases gives us confidence to consider the former more generally in higher d. In particular, a separate estimate of the fitting parameter  $\widehat{A}$  can be obtained from the HNC. Unlike the estimate from the virial series, this one is not affected by spurious poles, and hence its high-d behavior is smoother (see Fig. 3). The same overall trend is obtained. This combination of methods, thus, suggests that the asymptotic convergence to  $\widehat{A}^{CS}$  is rather slow (see the inset of Fig. 3).



**FIG. 3.** Fitting coefficients for the CS form,  $\widehat{A}^{\text{sim}}$  (purple solid line with stars), grow from a negative value for d < 5 to a peak in d = 9 before steadily decreasing. This behavior is consistent with the prediction for  $\widehat{A}^{\text{CS}31}$  (black line) to which the numerical results could plausibly converge at even higher d than is reached here. The quantity should, therefore, vanish as  $d \to \infty$ . The dimensional range is extended by fitting the virial (purple dashed line) and HNC (red line)<sup>32</sup> up to d = 50 to a CS form. High-d convergence is best assessed on a logarithmic scale (inset).

#### B. Pair distribution function

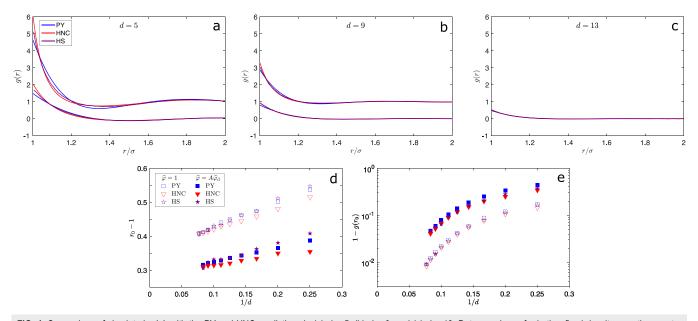
Although the equation of state is directly related to contact features of the pair correlation [see Eq. (3)], it provides no information about the liquid structure beyond that regime. In particular, it offers little insight into the first solvation shell, which includes the

particles that control self-caging as the density increases. Comparing the HNC and PY predictions with numerics in that range should therefore offer some insight into the quantitative reliability of  $\widehat{\varphi}_{\rm d}$  predictions.

Because g(r) depends on density, a proper dimensional comparison should be scaled accordingly. Here, because we are interested in the co-evolution of structure and dynamics, we specifically monitor the structure at two characteristic densities of sluggish liquids (further discussed in Sec. III):  $\widehat{\varphi}_{\rm nf} \approx 1$  and  $A\widehat{\varphi}_{\rm d}$  (the factor  $A\approx 0.93$  is chosen to produce  $\widehat{\varphi}$  which are close to  $\widehat{\varphi}_{\rm d}$ , yet remain computationally accessible for all d). Figures 4(a)–4(c) are consistent with our expectation that g(r) should deviate more strongly from a step function as the density increases but that these deviations should be systematically suppressed as d increases. Both the HNC and PY predictions then also appear to capture numerical results increasingly well.

To quantify this convergence more carefully, we specifically consider the position,  $r_1$ , and the depth,  $g(r_1)$ , of the first minimum of g(r), which traditionally delimit the end of the first solvation shell. Because g(r) tends toward a step function as  $d \to \infty$ ,  $1-g(r_1)$  should then vanish; however,  $r_1-1$  is not similarly constrained. Figure 4 shows that as  $d \to \infty$ , the former quantity indeed vanishes—and does so monotonically—and that the latter does not. The thickness of the first solvation shell, which in low d commonly describes the caging process, therefore does not asymptotically estimate the cage size, which vanishes instead as 1/d.

Interestingly, although predictions for both quantities from PY and HNC converge as  $d \to \infty$ , numerical simulation results generally appear to converge toward the HNC predictions more



**FIG. 4.** Comparison of simulated g(r) with the PY and HNC predictions in (a) d=5, (b) d=9, and (c) d=13. Data are shown for both a fixed density near the onset,  $\widehat{\varphi}=1$  (vertically offset by -1) and for a density near the simulation estimate for  $\widehat{\varphi}_d$  at  $A\widehat{\varphi}_d$  with  $A\approx 0.93$ . To quantify the visual differences between these descriptions, we consider (d) the position  $r_0$  and (e) the depth  $1-g(r_0)$  of the first minimum. In the  $d\to\infty$  limit, the depth and the position of the first minimum of HS, PY, and HNC all seem to converge. Note that because systematic deviations appear in d=11 and d=13 at  $\widehat{\varphi}=1$  for the system sizes considered in the rest of this work, they are here replaced with  $N\to\infty$  values (see the Appendix).

directly as *d* increases. The effect is marked and systematic for the minimum depth. The minimum position drifts from being above the PY prediction to near the HNC one as *d* increases, which here as well suggests that the low-*d* agreement with PY might be due to a fortuitous cancellation of error. Therefore, this analysis further supports the advantageous asymptotic dimensional scaling of the HNC in describing structural pair correlations in the liquid.

Given that computationally accessible system sizes for d = 13 at densities beyond the onset are markedly too small to capture the shell structure (see the Appendix), however, we do not consider these systems further.

#### IV. DIMENSIONAL EVOLUTION OF FLUID DYNAMICS

In this section, we compare the dimensional evolution of dynamical observables with theoretical predictions. Specifically, we consider the dynamical onset and  $\widehat{\varphi}_d$ .

# A. Dynamical onset

In order to correlate these structural observations with liquid dynamics, we first characterize the dynamical onset. To define this onset, we recall that at low densities, hard spheres that evolve under a MC (Brownian-like) dynamics are Fickian fluids with a MSD given by  $\langle r^2(t) \rangle = Dt$ , where D is a constant set by microscopic dynamics alone. In other words, the MSD is then featureless. As the density

increases, however, the fluid becomes non-Fickian, thus hinting at the emergence of non-trivial dynamics. This transformation, however, is a crossover even in mean-field descriptions, and hence, it is not uniquely defined. We here follow the convention of Refs. 8 and 9 in identifying the dynamical onset from the emergence of an inflection point in the (logarithmically scaled) MSD (see Fig. 1a),

$$\frac{\mathrm{d}^2(\ln\langle r^2(t)\rangle)}{\mathrm{d}(\ln t)^2} = 0. \tag{11}$$

The resulting estimate is consistent with  $\widehat{\varphi}_{nf} = 1.3(1)$  in all d, which agrees remarkably well with molecular dynamics results8 (see Table II). The choice of microscopic dynamics, therefore, appears to have a relatively weak effect on that onset, at least in the d range studied. By contrast, recent work by Manacorda et al. finds an onset slightly below unity for both Newtonian and Brownian dynamics for  $d \to \infty^9$  in agreement with the numerical results for minimally structured yet finite-d (Mari–Kurchan) models. It, therefore, seems natural to conclude that the structure of HS fluids underlies the ~ 30% difference between finite-d and  $d \rightarrow \infty$  results. However, What particular structural feature could explain this difference? The shell structure of low-d HS fluids may be responsible, but one would then (naively) expect the dynamical onset to emerge at a *lower*  $\widehat{\varphi}$  than what the  $d \to \infty$  solution predicts. The difference between finite-d and  $d \to \infty$  results is nevertheless rather small, especially relative to the marked increase in  $\widehat{\varphi}_d$  with d (see Sec. IV B). Both the structural

**TABLE II.** Pseudo-critical parameters  $(\widehat{\varphi}_d, \widehat{p}_d, \text{and } \gamma)$  as well as the dynamical onset density  $\widehat{\varphi}_{nf}$  from this work (bold) along with the estimates from Refs. 8, 13, and 20. The fit parameter  $\widehat{A}$  to the CS form in Eq. (8) is also included for reference.

d	$\widehat{arphi}_{ m d}$	$\widehat{p}_{ m d}$	$\widehat{oldsymbol{arphi}}_{ m nf}$	γ	$\widehat{A}$	Reference
4	1.624			2.4(3)		20
	1.60(2)					13
	1.6144(8)		1.17(2)	1.92(3)		8
	1.597(7)	6.11(10)	1.30(2)	1.70(11)	-0.0278(17)	
5	1.71(3)					13
	1.7171(6)		1.22(13)	1.95(3)		8
	1.701(5)	5.34(5)	1.36(8)	1.82(8)	0.1704(19)	
6	1.83(5)					13
	1.8379(11)		1.17(11)	2.00(3)		8
	1.827(5)	4.88(4)	1.33(5)	1.94(8)	0.367(3)	
7	1.94(9)					13
	1.968(2)		1.19(9)	2.0(1)		8
	1.956(12)	4.46(6)	1.29(10)	2.00(14)	0.511(2)	
8	2.07(2)					13
	2.107(2)		1.28(6)	2.15(5)		8
	2.090(11)	4.17(4)	1.36(8)	2.05(10)	0.5994(19)	
9	2.19(3)					13
	2.222(5)	3.901(17)	1.28(3)	2.03(4)	0.6228(18)	
10	2.31(5)					13
	2.364(5)	3.76(2)	1.30(8)	2.09(4)	0.613(3)	
11	2.44(9)					13
	2.49(2)	3.56(5)	1.28(4)	2.09(10)	0.567(5)	
12	2.6(1)					13
	2.60(4)	3.43(8)	1.31(10)	2.02(19)	0.522(5)	

origin and the meek dimensional dependence of  $\widehat{\phi}_{nf}$  are, therefore, puzzling. Without a crisper understanding of the finite-d mean-field dynamics, however, further insight likely remains out of reach.

# **B.** Dynamical transition

We finally consider the dynamical transition regime. At a genuine dynamical transition, perfect caging would result in the MSD exhibiting an infinitely extended plateau. In addition, upon approaching the transition density, the diffusivity and the structural relaxation time would diverge, scaling critically as

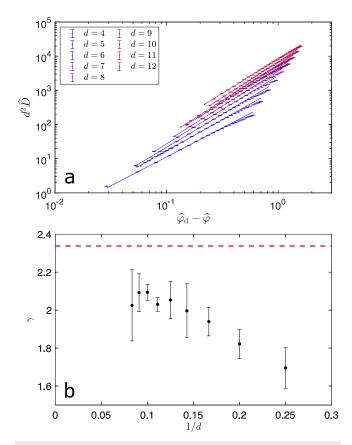
$$D^{-1} \sim \tau \sim (\varphi_{\rm d} - \varphi)^{-\gamma}. \tag{12}$$

Like any spinodal critical point, however, a true dynamical transition can only exist in the mean-field limit of high-spatial dimension,  $d \to \infty$ . In finite-d systems, activated processes blur the singularity into a crossover and decouple D and  $\tau$ , thus giving rise to a breakdown of the Stokes–Einstein relation. If the activated processes can be screened somehow or if the dimension is sufficiently high, traces of criticality may nevertheless be distinguished, including the pseudo-critical scaling of Eq. (12). Because  $\tau$  naturally screens some of the activated processes,  $\tau$  its scaling is commonly used to explore this regime.

A reliable estimator of  $\tau$ , however, requires knowing the cage size, as parameterized by a in Eq. (1). In low d, this quantity is traditionally estimated from the first peak of the structure factor, 48 but as d increases, the typical cage size grows increasingly disconnected from this structural feature.<sup>14</sup> [Because the width of the first peak of g(r) does not vanish as  $d \to \infty$ , as discussed in Sec. III B, this quantity is also a poor estimator of the cage size as d increases]. The proper a must instead be determined from the cage size near  $\varphi_d$ , such as by taking the square root of the MSD plateau height. Because we here have set a = 0.3, which is much larger than the cage size as d increases, the associated characteristic relaxation times systematically overestimate  $\tau$  and the pseudo-critical scaling of these relaxation times leads to unphysically large  $\gamma$ . We, thus, leave a detailed consideration of the scaling of  $\tau$  to future studies and here determine  $\gamma$  soley from the scaling of D. In practice, its value is extracted by fitting the long-time scaling of the MSD using the empirical form

$$\langle r^2(t) \rangle / t = D + a_0 t^{-a_1},$$
 (13)

where fit parameters  $a_0, a_1 > 0$  depend on  $\varphi$ . (Their specific values are unimportant for the subsequent analysis.) However, this approach is not completely devoid of difficulties. For MC dynamics, in particular, at microscopic times  $t \lesssim \mathcal{O}(1)$ , the system behavior is roughly independent of density. At asymptotically small D, this effect is negligible, but given the limited available dynamical regime, this effect cannot be brushed aside. In order to calibrate our results, we, thus, use of a multiplicative scale factor that fixes the MSD at t=1 for all densities considered. In addition, the pseudo-critical scaling regime of Eq. (12) can only be expected to hold between  $\varphi_{\rm nf}$  and the density at which the Stokes–Einstein relationship (SER) is found to break  $\varphi_{\rm SER}$  (see Ref. 8). Even this range is too broad. The  $d \to \infty$  scaling of D deviates markedly from the asymptotic power-law scaling already at  $\widehat{\varphi} \sim \widehat{\varphi}_d/2$ . Using this observation as a guide, we here only fit the upper half of the density regime beyond



**FIG. 5.** The (pseudo-)critical scaling of the scaled diffusivity  $\widehat{D}$  from Eq. (12) (offset vertically by a factor of  $d^2$  for clarity) over the relevant density regime (see the text) provides estimates for both (a)  $\widehat{\varphi}_d$  and (b) the associated critical exponent,  $\gamma$ . For  $d \geq 4$ ,  $\gamma$  rises monotonically and approaches the  $d \to \infty$  value of  $\gamma = 2.337\,86^7$  (red dashed line). Numerical values of  $\widehat{\varphi}_d$  and  $\gamma$  are reported in Table II.

the onset of non-Fickian diffusion to the (pseudo-)critical scaling form of Eq. (12) while staying clear of the SER breakdown regime in low d (see Fig. 5). (In practice, we use a lower bound such that  $\widehat{\varphi} > (1 + \widehat{\varphi}_d)/2$ .)

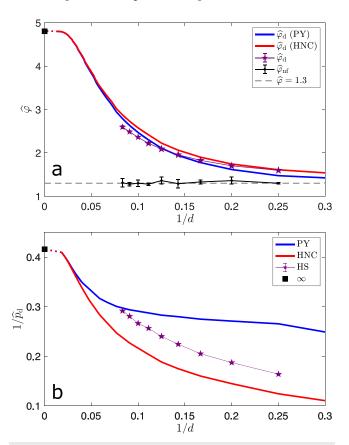
The resulting estimates of  $\widehat{\varphi}_d$  and of the non-universal critical exponent  $\gamma$  can be seen in Table II and Fig. 5. The resulting  $\widehat{\varphi}_d$  are marginally different but generally consistent. They also are consistent with previously reported values. The results for  $\gamma$  obtained via Monte Carlo dynamics also agree, within error bars, with the molecular dynamics estimates and extend by nearly 50% the dimensional range of under solid numerical control.

The larger error bars compared to those of Refs. 8 and 13. reflect the increased understanding of systematic errors involved in simulating glass forming liquids and extracting the pseudo-critical behavior. We, thus, expect the current estimates to hold for the foreseeable future. The simultaneous fit of  $\widehat{\varphi}_d$  and y, however, leaves some undue wiggle room in the analysis. Because the resulting power-law scaling depends fairly sensitively on  $\widehat{\varphi}_d$ , which itself increases fairly rapidly with d, the finite-d theoretical estimates for the quantity are unlikely to ever serve as efficient substitutes. By contrast,  $\gamma$  increases slowly and monotonously with d toward its  $d \to \infty$ 

value,  $\gamma = 2.337 \, 86.^7$  If a finite-d estimate of that exponent was independently determined (using, e.g., the approach of Ref. 49), then a more controlled test of the theory and a more reproducible estimate of  $\widehat{\varphi}_{\rm d}$  should be possible.

Mangeat and Zamponi<sup>12</sup> have shown that both the HNC and PY predictions provide correct order-of-magnitude estimates for  $\widehat{\varphi}_d$  in low d and converge with each other for  $d \gtrsim 30$ . Simulation results, however, show that while both the HNC and PY approximations capture the general trend of  $\widehat{\varphi}_d$  in finite d, both overshoot it beyond even our enlarged error bars [Fig. 6(a)] in the highest d attained. One potential source of discrepancy is the reliance of Mangeat and Zamponi on the Gaussian cage approximation, which breaks down in finite-d hard spheres<sup>8,43</sup> and is known to lead to significant 1/d corrections in a related model.<sup>50–52</sup> The dimensional range accessible in simulations is, however, too small to assess this hypothesis. Structural deviations at lower d are indeed too pronounced for their dynamical contribution to be properly ascertained.

In any event, considering the success of HNC in predicting the equation of state and the pair structure, our results are consistent with the suspicion of Mangeat and Zamponi that the HNC estimates



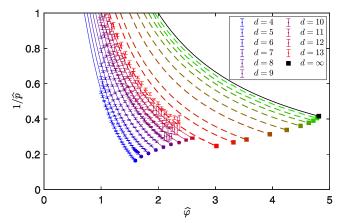
**FIG. 6.** Dimensional evolution of (a)  $\widehat{\varphi}_d$  and the dynamical onset  $\widehat{\varphi}_{nf}$  as well as (b)  $\widehat{\rho}_d$ . The onset should be compared with the mean-field prediction  $\widehat{\varphi}_{nf} \leq 1$ . The predicted mean-field values for  $\widehat{\varphi}_d$  using both PY and HNC  $^{12}$  deviate from the numerical estimates, but the HNC approximation describes their trend more reliably. Note that the PY and HNC results for  $\widehat{\rho}_d$  and  $\widehat{\varphi}_d$  can only be obtained up to d=55 and d=70, respectively, using the available codes. A dashed line is then used to connect these with the  $d\to\infty$  results.

for  $\widehat{\varphi}_d$  (and the corresponding pressure  $\widehat{p}_d$ ) are at best an upper bound. Similarly, the inadequacy of PY in capturing the equation of state and the pair structure at intermediate d results in predictions for  $\widehat{\varphi}_d$  that overshoot and for  $\widehat{p}_d$  that undershoots the numerical results.

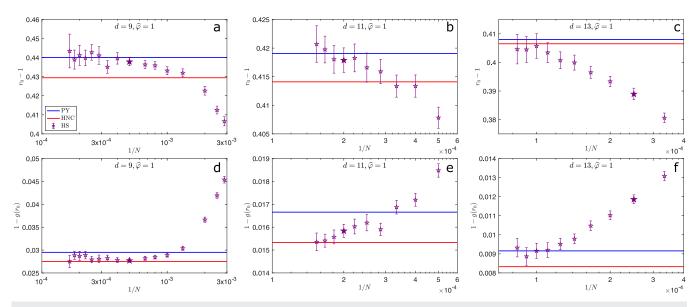
#### V. CONCLUSION

Our results suggest that the gap between the  $d \to \infty$  liquid structure and numerical simulations is essentially closed. As the dimension increases, the HNC tracks the smoothing of the pair correlation function and nearly quantitatively captures simulation and virial results for  $d \gtrsim 13$ . (PY, however, fails to capture the key structural features in intermediate dimensions, d = 4-13.) Given this match, we also understand that (i) non-trivial structural corrections are almost irrelevant for  $d \gtrsim 50$  and that (ii) higher-order correlations are necessary to quantitatively describe systems with d < 13.

Our results, however, also suggest that a certain gap in the dynamical description persists (see Fig. 7). Existing mean-field theory-based predictions for the dynamical onset and the dynamical transition are close to the simulation estimates but remain quantitatively distinct (even when the pair structure description is nearly flawless). Can more be done to bridge the disconnect? Unfortunately, numerical techniques seem to have approached their practical limit. Properly simulating higher d fluids would require either an enormous computational undertaking or some marked innovation in numerical techniques. In the near term, theoretical improvements are more likely. In particular, a first-principles calculation for  $\gamma$  should be within reach, and advances on the dynamical mean-field theory might eventually provide some quantitative insight. It is also possible that subtle higher-order correlations could also impact the mean-field-like dynamics, as has recently been suggested,  $^{54}$  but



**FIG. 7.** Liquid equations of state for d=4-12 (graded from blue to red) obtained by fitting simulation results (symbols) to the generalized CS form in Eq. (8) along with the  $d\to\infty$  equation of state from Eq. (4) (black line). Squares denote  $(\widehat{\varphi}_d,\widehat{\rho}_d)$  obtained as in Sec. III, and  $\widehat{\varphi}=4.81\ldots$  for  $d\to\infty$ . The HNC equations of state for d=15,17,21,25,30,35,40, and 50 (dashed lines graded from red to green) are depicted up to the predicted mean-field dynamical transition. In d=13, simulation pressure data are shown, but the CS fit is omitted, and the HNC equation of state is shown instead (red dashed line).



**FIG. 8.** Finite-size scaling of the position of the first minimum of the pair correlation function  $r_0 - 1$  in (a) d = 9, (b) d = 11, and (c) d = 13 as well as the value that the pair correlation takes in (d) d = 9, (e) d = 11, and (f) d = 13, all taken at  $\widehat{\varphi} = 1$ . In all dimensions, the system size used in this work is denoted with a filled in symbol. The results quickly converge as N increases. The difference from the thermodynamic limit is indistinguishable in d = 9, is quite small in d = 11, but is significant in d = 13. The PY and HNC predictions are included to facilitate comparison with Fig. 4.

this proposal cannot be properly assessed without a more robust dynamical theory.

#### **ACKNOWLEDGMENTS**

We thank Nathan Clisby for useful discussions about the virial expansion and Francesco Zamponi for various discussions. We would also like to thank Atsushi Ikeda for sharing a code that solves the HNC structure and Andres Santos for a code that solves the PY structure. This work was supported by a grant from the Simons Foundation (Grant No. 454937 to P.C.). The simulations were performed both at Duke Compute Cluster (DCC)—for which the authors thank Tom Milledge's assistance—and on Extreme Science and Engineering Discovery Environment (XSEDE), which is supported by the National Science Foundation (Grant No. ACI-1548562).

# **AUTHOR DECLARATIONS**

#### **Conflict of Interest**

The authors have no conflicts to disclose.

#### **DATA AVAILABILITY**

The data that support the findings of this study are openly available in Duke Digital Repository.<sup>55</sup>

# APPENDIX: FINITE-SIZE SCALING IN d = 9-13

As mentioned in the main text, the structure and dynamics of supercooled liquids generally depend only weakly on the system size,

provided that the simulation box is large enough to contain a representative environment for a given particle. Equation (7), however, shows that the system size necessary to overcome this problem grows exponentially with d. Simulation results are thus most sensitive to the system size in the highest d considered, which approach the limit of current computational feasibility.

Figure 8 shows how the first minimum of g(r) at the onset density evolves with N (as in Fig. 4) for d = 9-13. In all cases, the system size used in this work is noted. Upon increasing d, the structural observables appear to plateau at larger N than what is used, indicating that significant corrections to the structure (and hence dynamics) then persist. For d = 13, the discrepancy is quite notable, but for d = 11, it is relatively small, and for d = 9, it is non-existent. The marked system-size artifacts on the liquid structure in d = 13 lead us to exclude that dimension from the subsequent dynamical analysis.

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